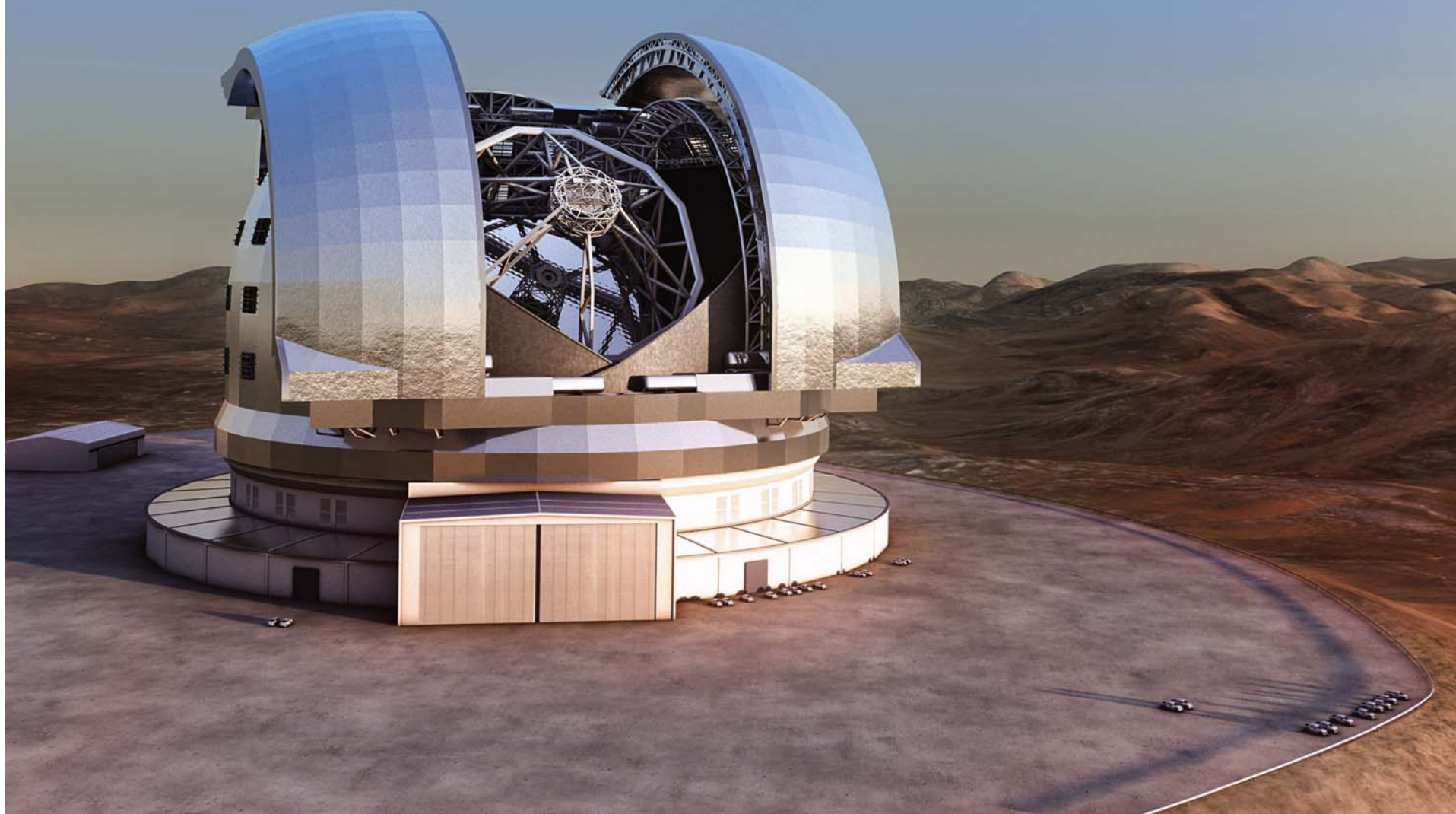


EUROPEAN EXTREMELY LARGE TELESCOPE

RAMBOLL



BUDGET: 8 MIA DKK



ASTRONOMY

“The study of everything beyond the Earth”

Objects are far away, hence appear small and faint:

- Need for large telescopes: resolution and sensitivity

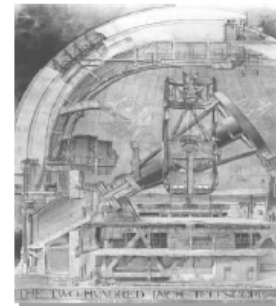
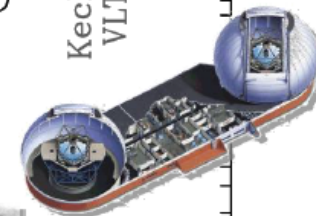
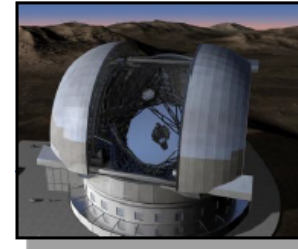
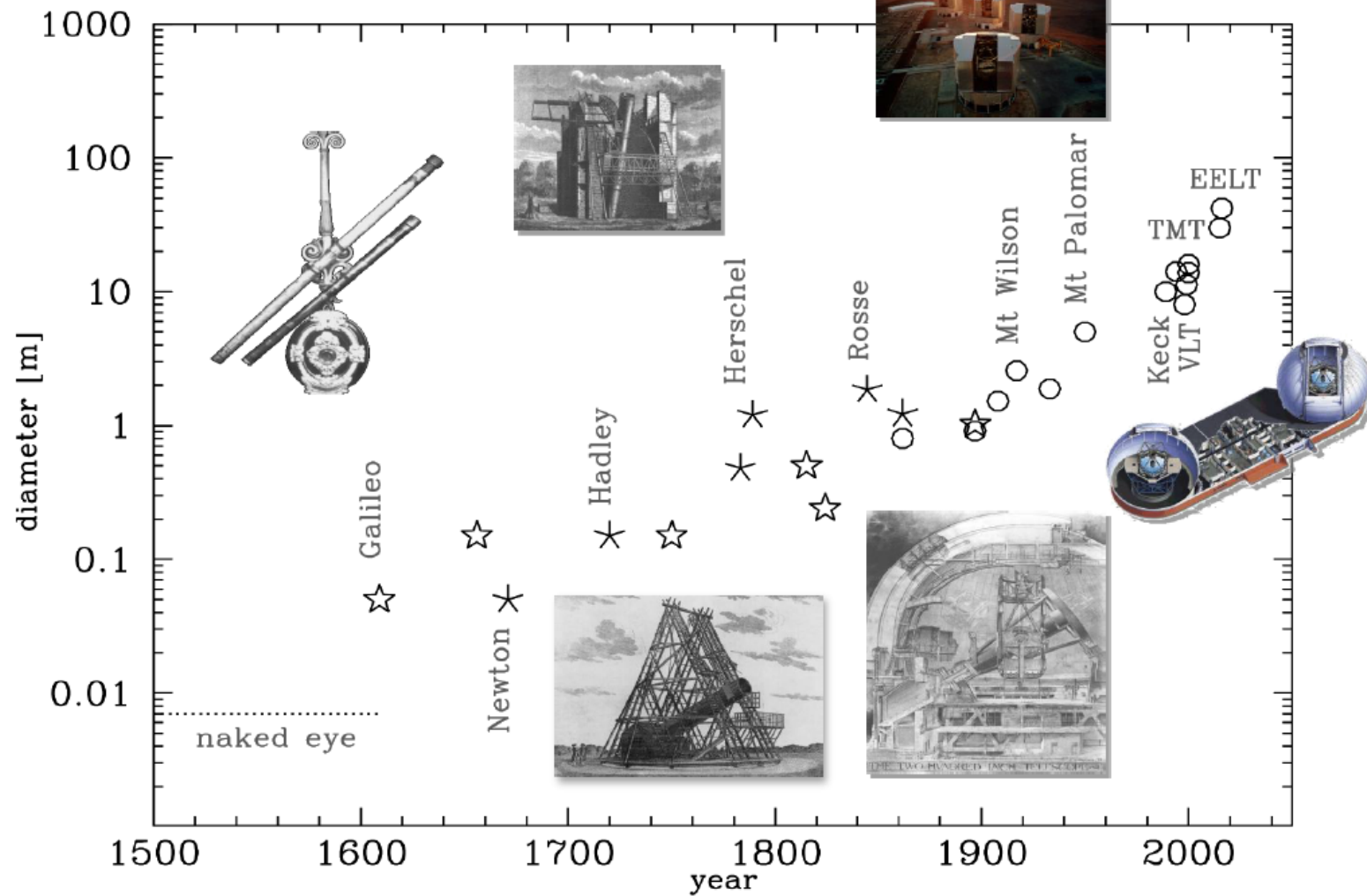
Observationally-driven science:

- Visual light and radio signals can be detected with telescopes on the ground

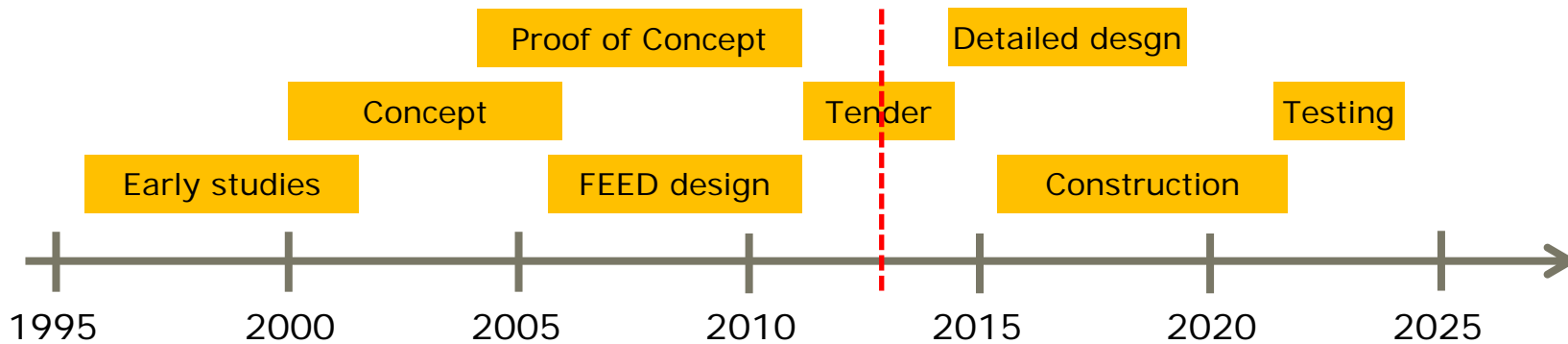
Technology now available to:

- Study objects over 95% of the age of the Universe
- Detect and study planets around other stars

TECHNOLOGY IN ASTRONOMY



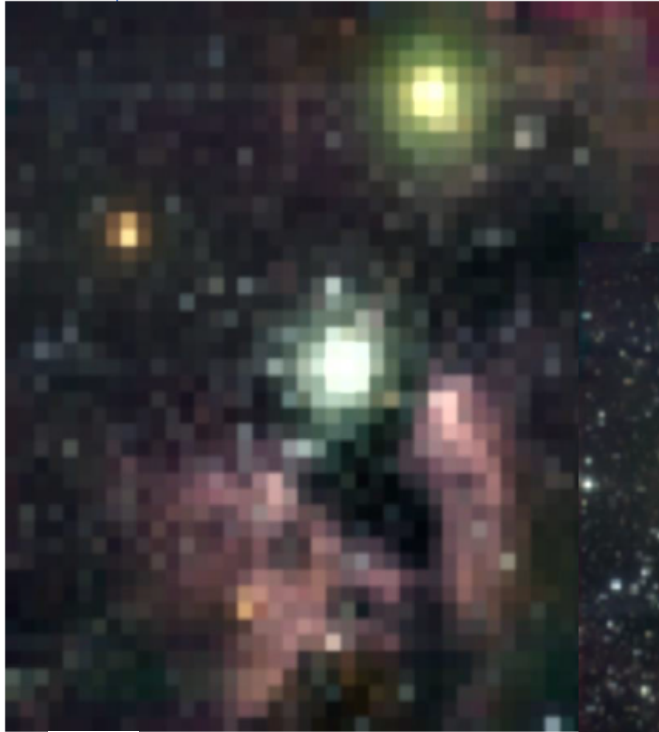
THE HISTORY OF ELT



- Scientist from all over Europe collaborate
- The ELT project has already been going on for 15 years
- In the beginning nobody knew what was possible
- Ramboll helped design the OWL* already in year 2000

* *Overwhelmingly Large Telescope*

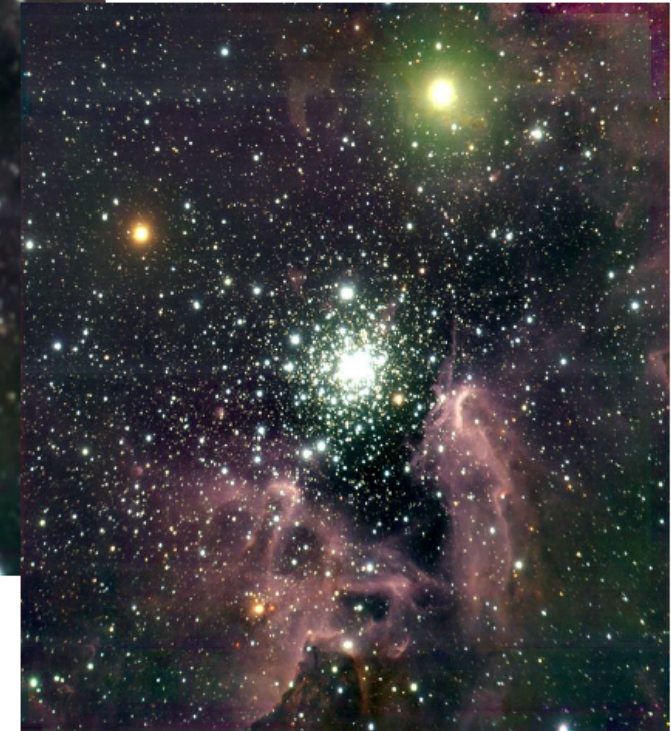
SHARPER PICTURES = SEE MORE



HST



VLT+AO



E-ELT

HST = Hubble Space Telescope
VLT = Very Large Telescope
ELT = Extremely Large Telescope
AO = Active Optics

ELT COMPARED TO HST



ESO – EUROPEAN SOUTHERN OBSERVATORY

1962

- ESO created by five Member States with the goal to build a large telescope in the southern hemisphere
- This became the 3.6m telescope on La Silla (1976)
- Denmark joined in 1967

2013

- 14+1 Member States (~30% of the world's astronomers)
- Paranal is the world-leading ground-based observatory
- ALMA (in partnership) on Chajnantor completed in 2013
- Construction of 39m E-ELT on Armazones to start soon

ESO LOCATIONS

La Silla

- First site, 3.6 m telescope

Chajnantor

- 5100 m altitude, radio waves

Paranal

- 3000 m altitude, VLT

Armazones

- 3000 m altitude, ELT

RAMBOLL



EUROPEAN VERY LARGE TELESCOPE ON PARANAL



EUROPEAN VERY LARGE TELESCOPE ON PARANAL



ALMA & APEX ON CHAJNANTOR

Purpose to detect submillimeter radio waves from the Universe
Requires minimal water vapour: 5100m altitude

APEX:

- 12m antenna in partnership with Sweden and Germany

ALMA

- 66 antennas: tremendous resolution and sensitivity
- Global partnership with North America and East Asia



EUROPEAN EXTREMELY LARGE TELESCOPE ON AMAZONES

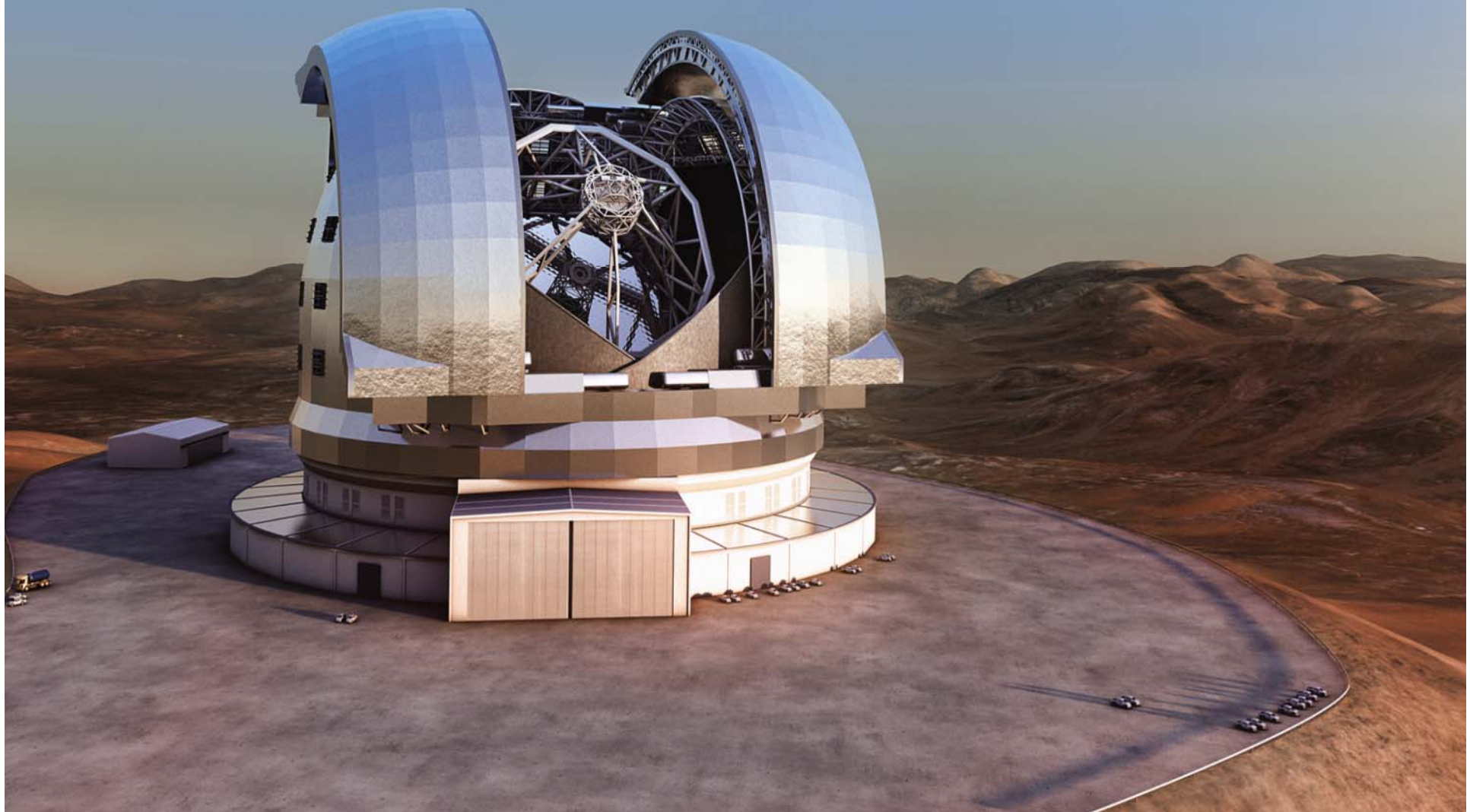
Largest optical/infrared telescope in the world

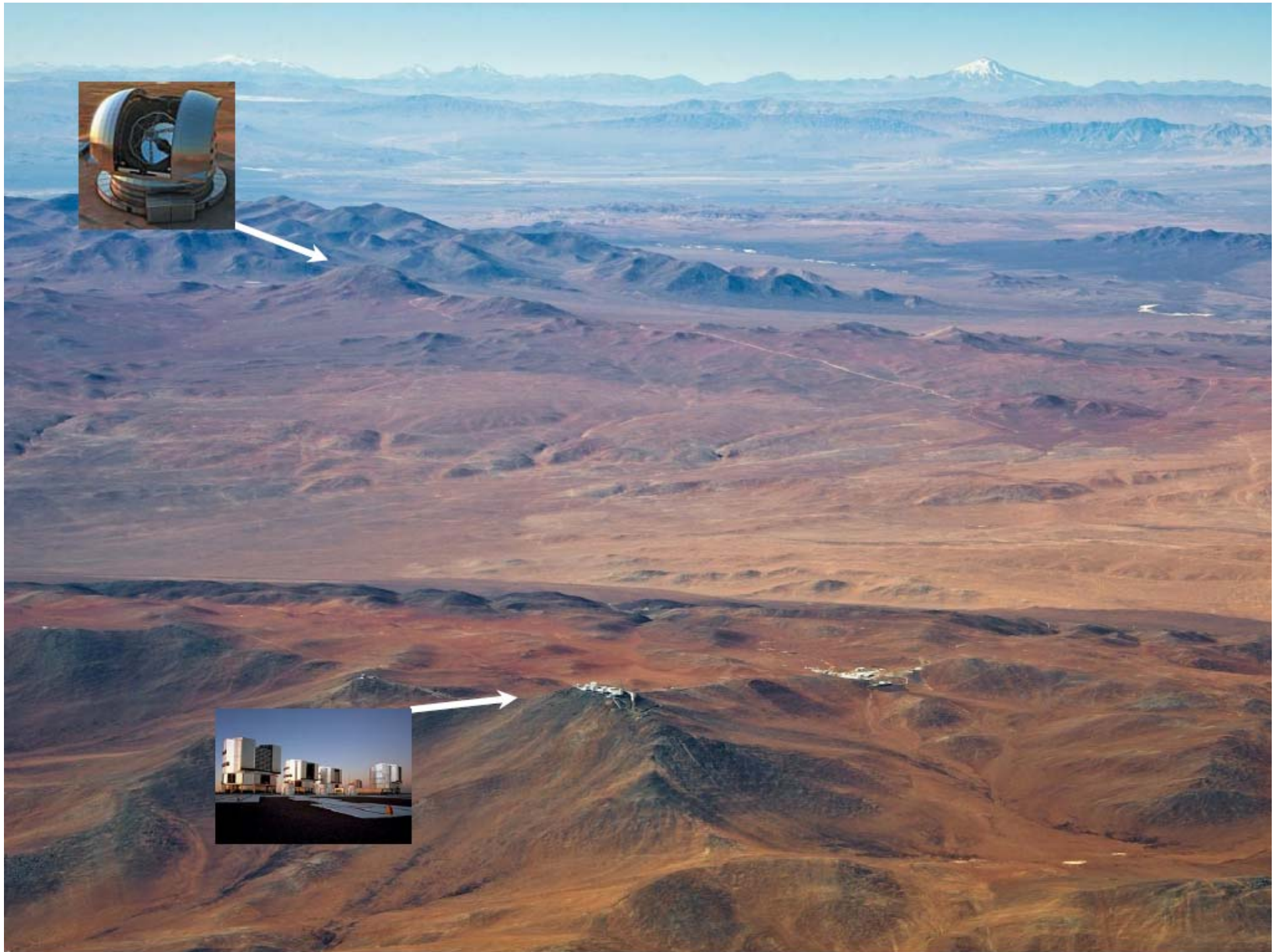
- 39m segmented primary mirror: transformational step
- Front end design complete, incl. instrumentation roadmap

Project

- Construction 2013-2024, on Cerro Armazones
- As integral part of the Paranal Observatory
- ESO cost: ~1100 MEUR incl. instruments and contingency

EUROPEAN EXTREMELY LARGE TELESCOPE ON AMAZONES





Five-mirror design

1. The 39.3-metre primary mirror collects light from the night sky and reflects it to a smaller mirror located above it.
2. The 4-metre secondary mirror reflects light back down to a smaller mirror nestled in the primary mirror.
3. The third mirror relays light to an adaptive flat mirror directly above.
4. The adaptive mirror adjusts its shape a thousand times a second to correct for distortions caused by atmospheric turbulence.
5. A fifth mirror, mounted on a fast-moving stage, stabilises the image and sends the light to cameras and other instruments on the stationary platform.

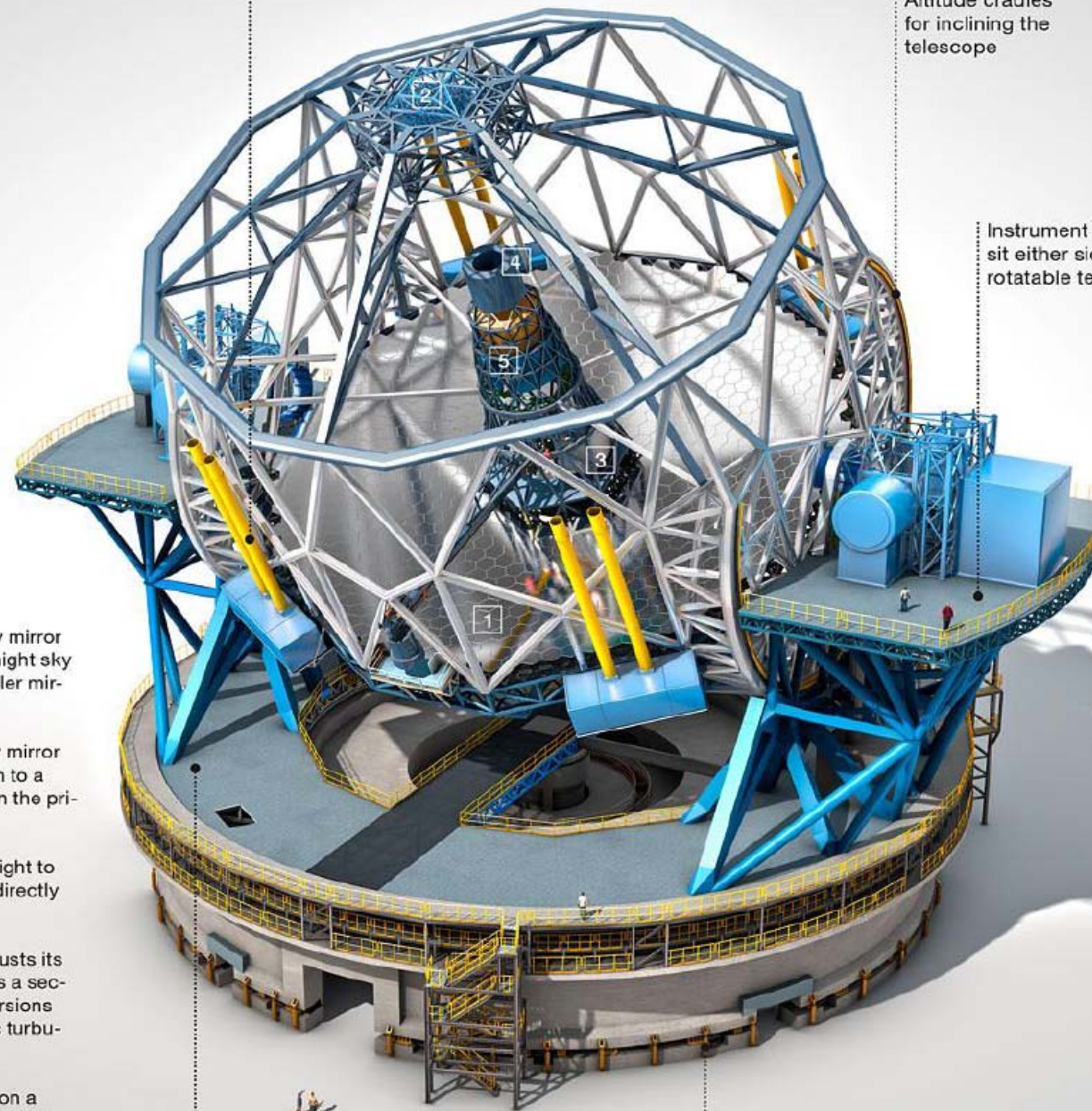
Lasers

Altitude cradles for inclining the telescope

Instrument platforms sit either side of the rotatable telescope

The 2800-tonne telescope system can turn through 360 degrees

Seismic isolators

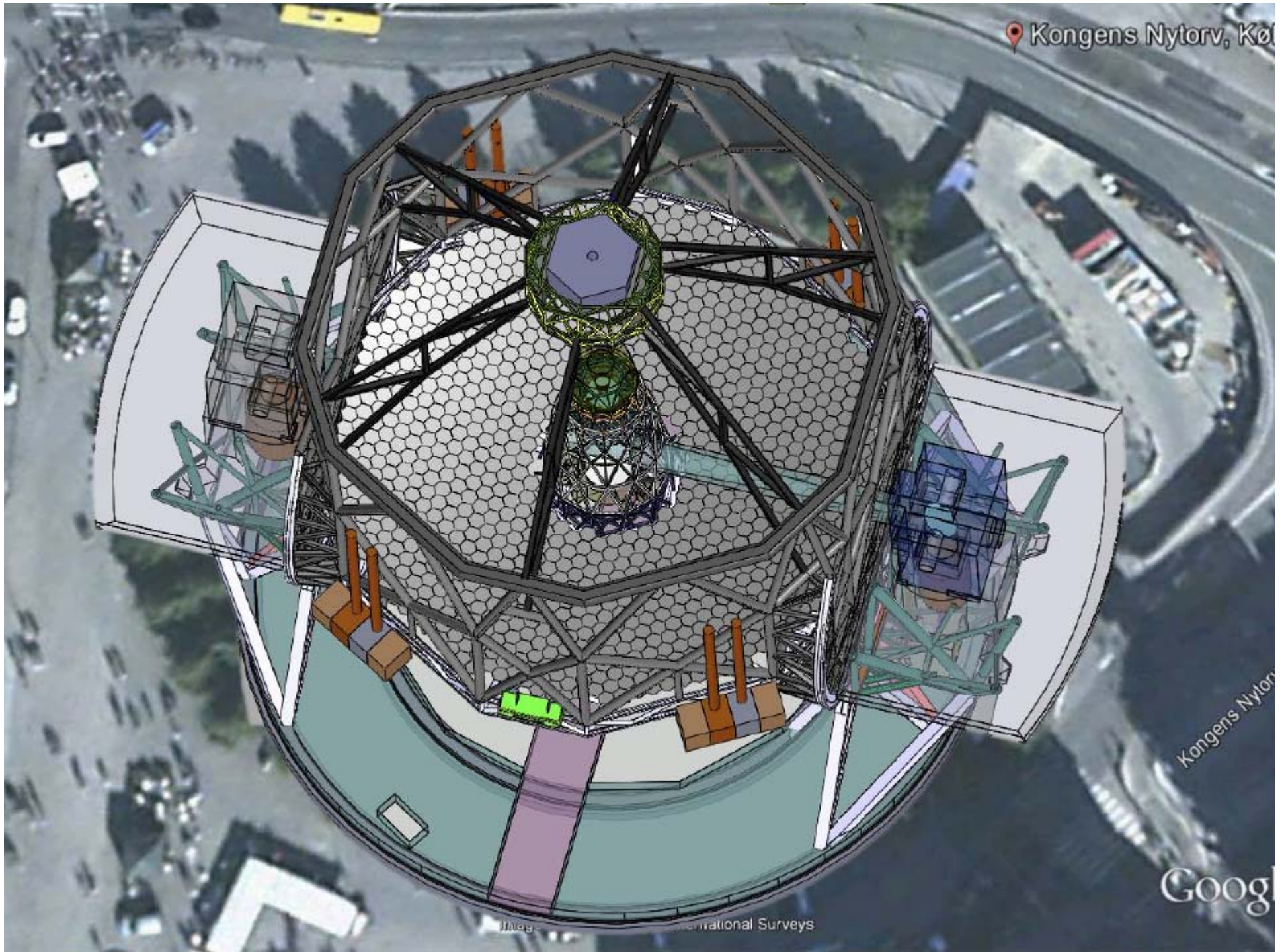


Kongens Nytorv, Kø

Kongens Nytorv

Google

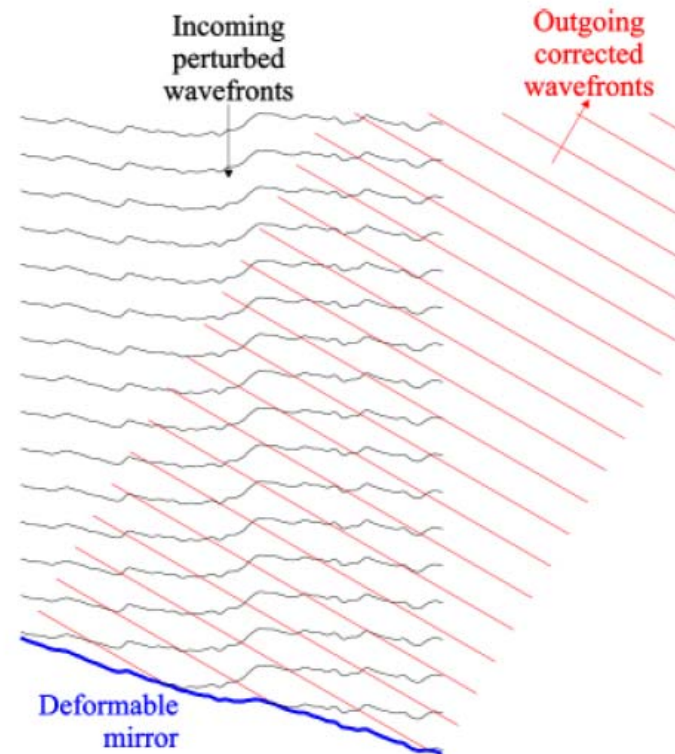
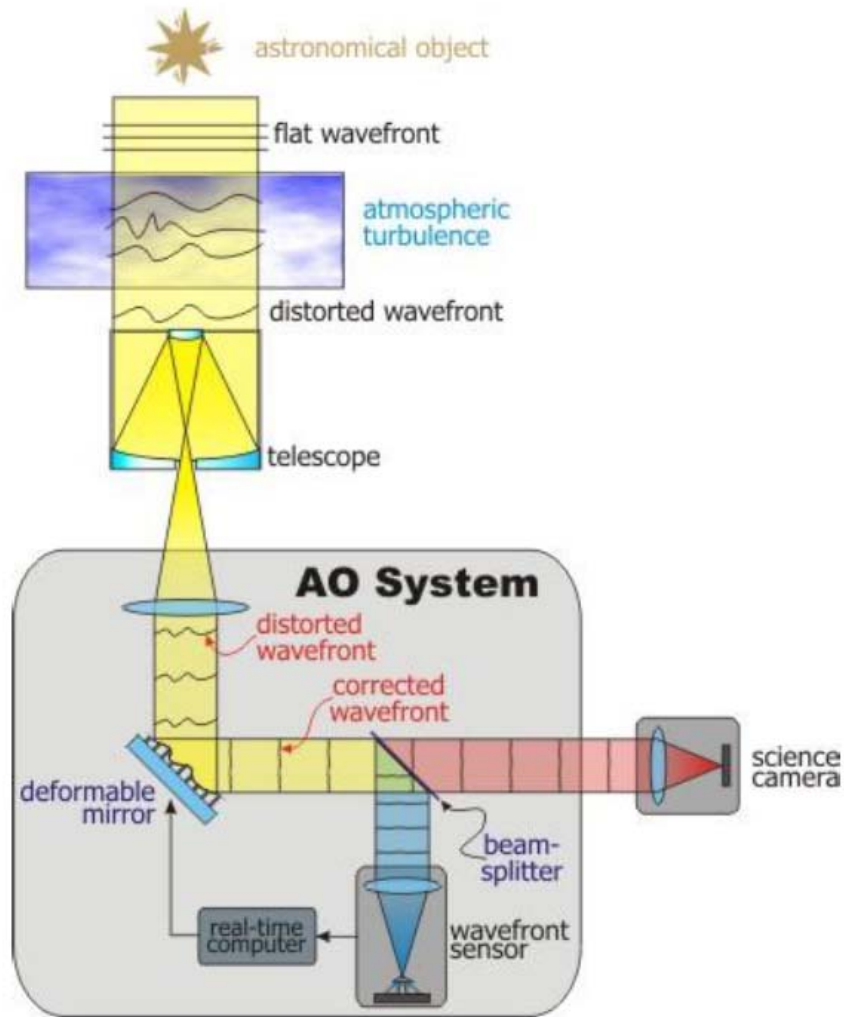
International Surveys



NEW TECHNOLOGIES DEVELOPED BY ESO

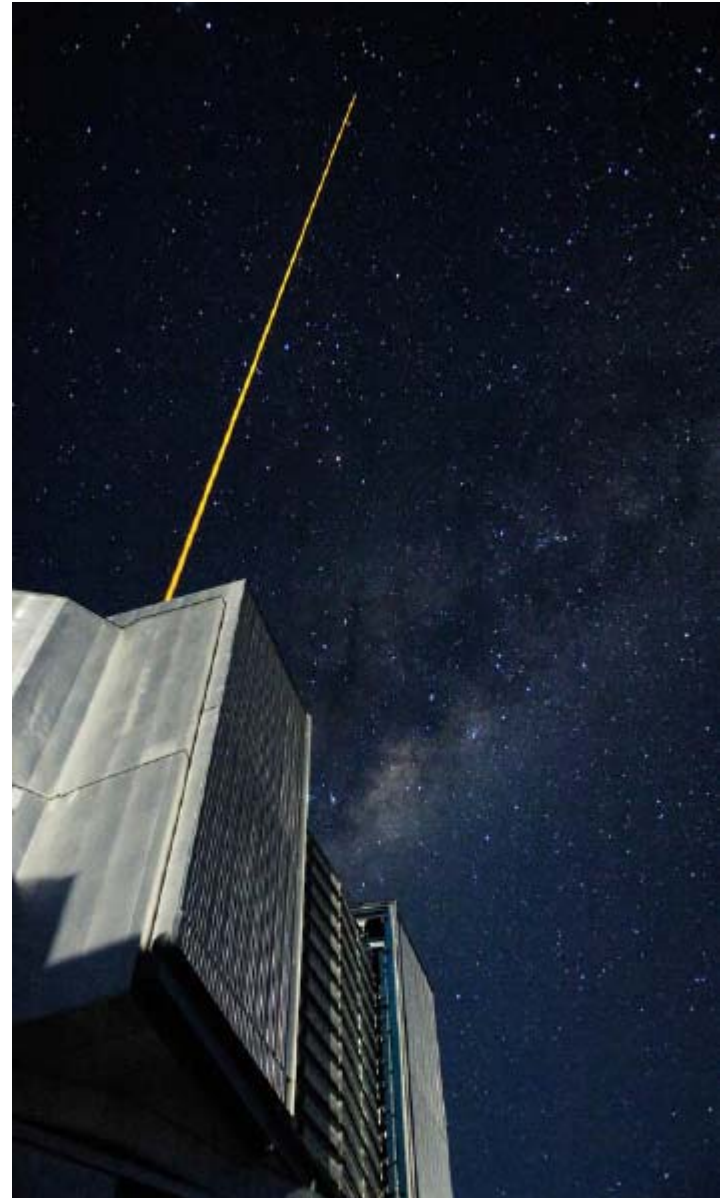
- ACTIVE OPTICS, now in use in most modern medium and large telescopes. It preserves optimal image quality by pairing a “flexible” mirror with actuators that actively adjust the mirror's shape during observations.
- ADAPTIVE OPTICS, the bigger a mirror, the greater its theoretical resolution, but even at the best sites for astronomy, large, ground-based telescopes observing at visible wavelengths cannot achieve an image sharpness better than telescopes with a 20- to 40-cm diameter, due to distortions introduced by atmospheric turbulence. One of the principal reasons for launching the Hubble Space Telescope was to avoid this image smearing.

ADAPTIVE MIRROR TECHNOLOGY



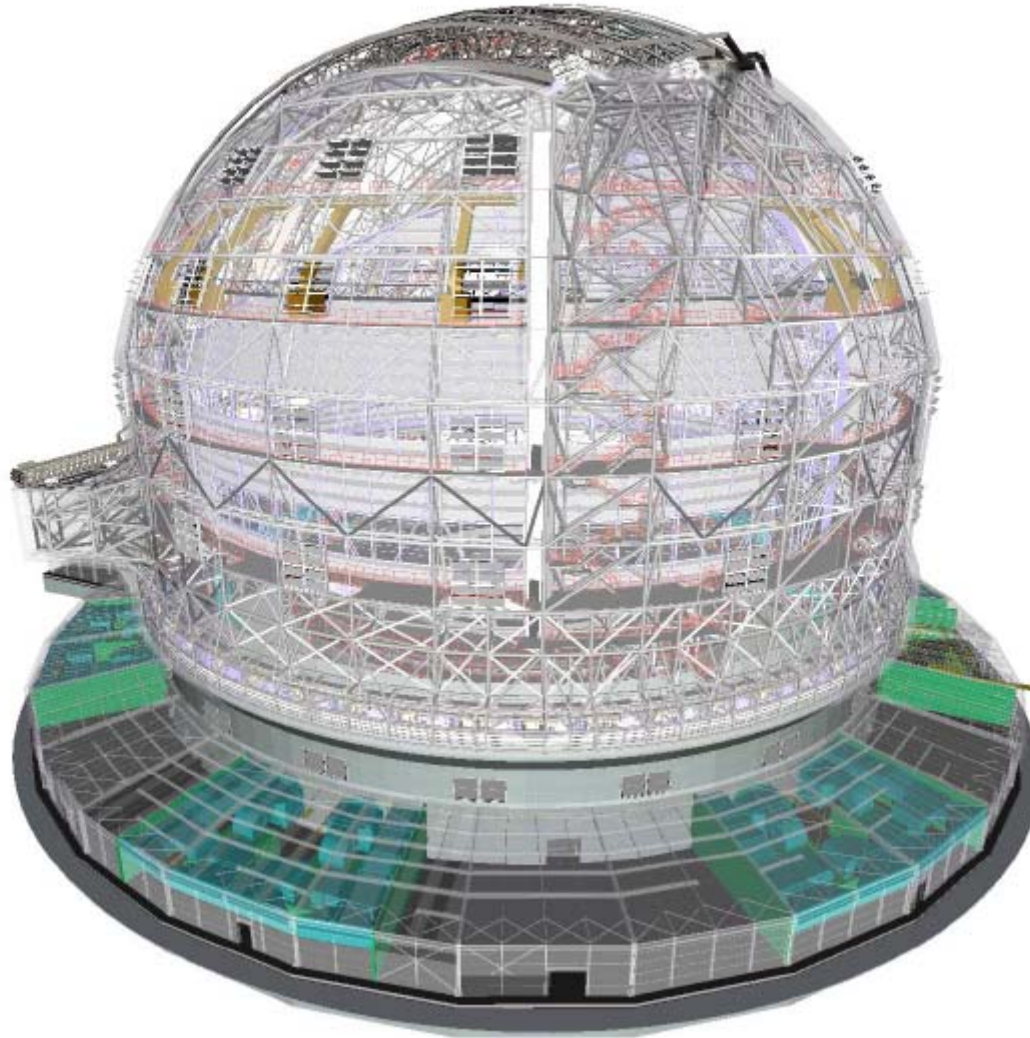
LASER GUIDE STARS

- Laser guide stars are artificial stars generated by exciting atomic sodium in the mesosphere at a height of 90km
- This requires a powerful laser beam launched from the telescope
- The yellow wavelength (589nm) is the colour of a sodium street lamp

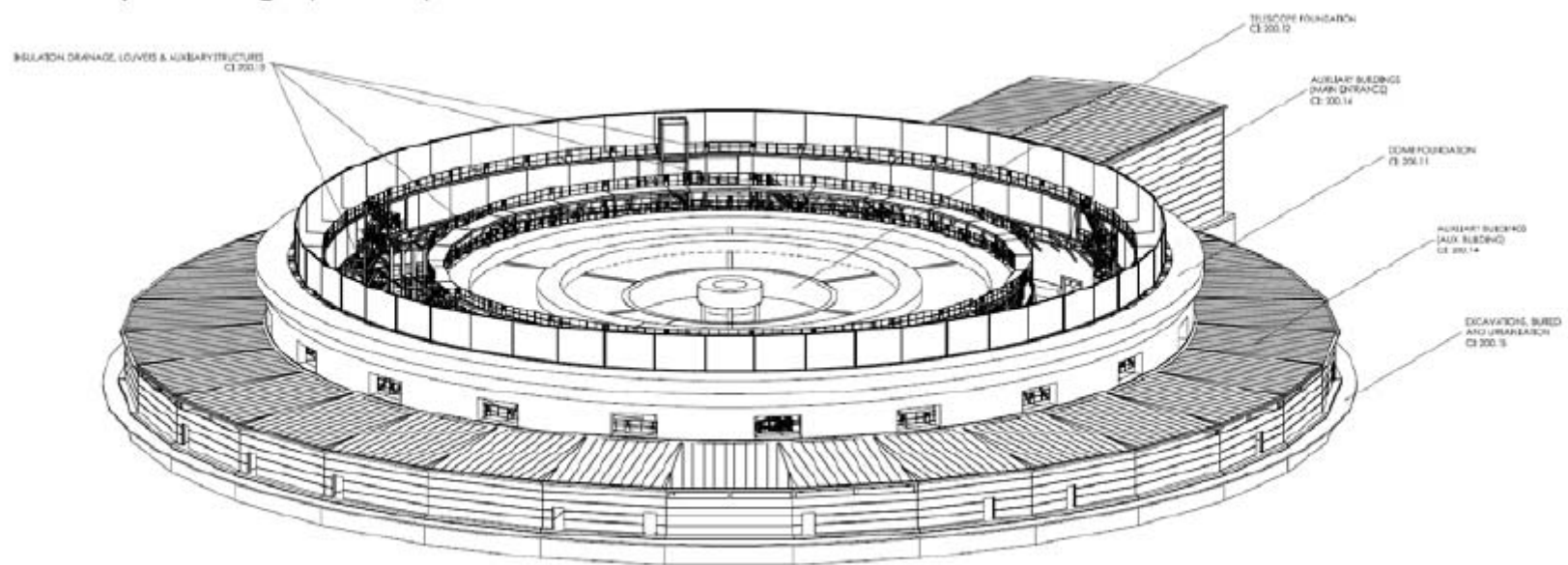




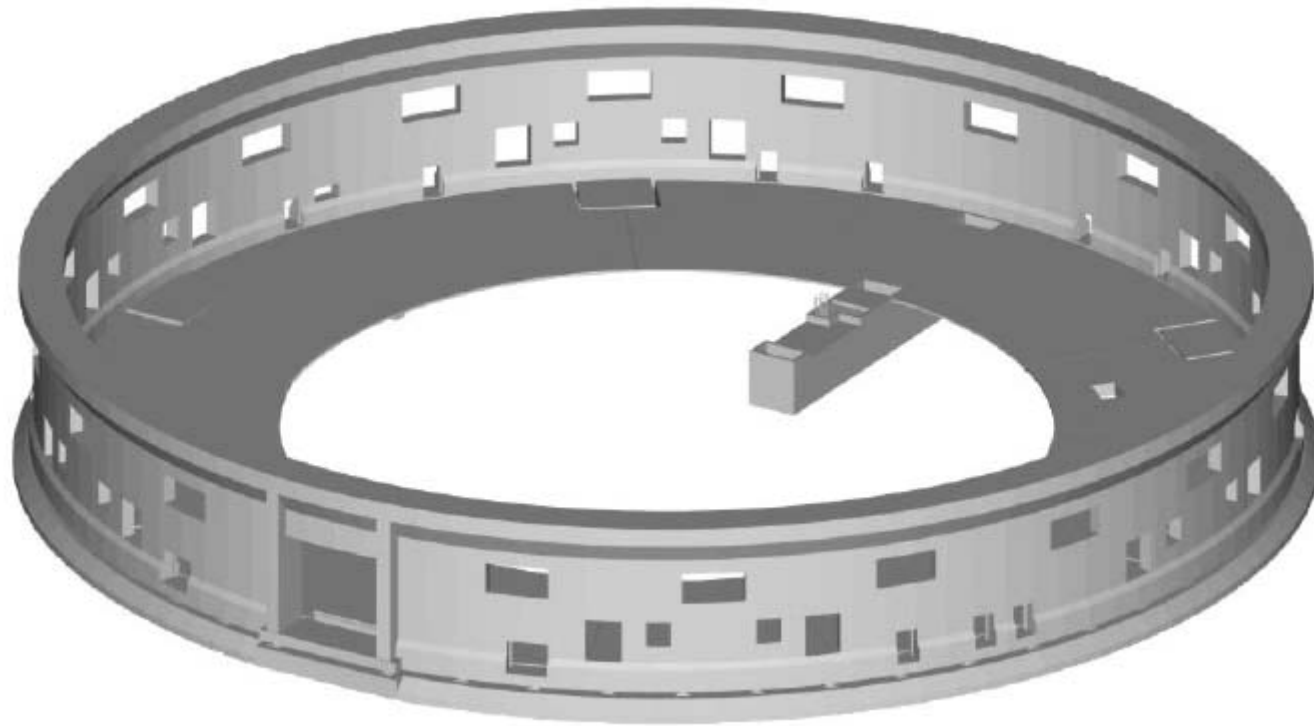
STRUCTURES



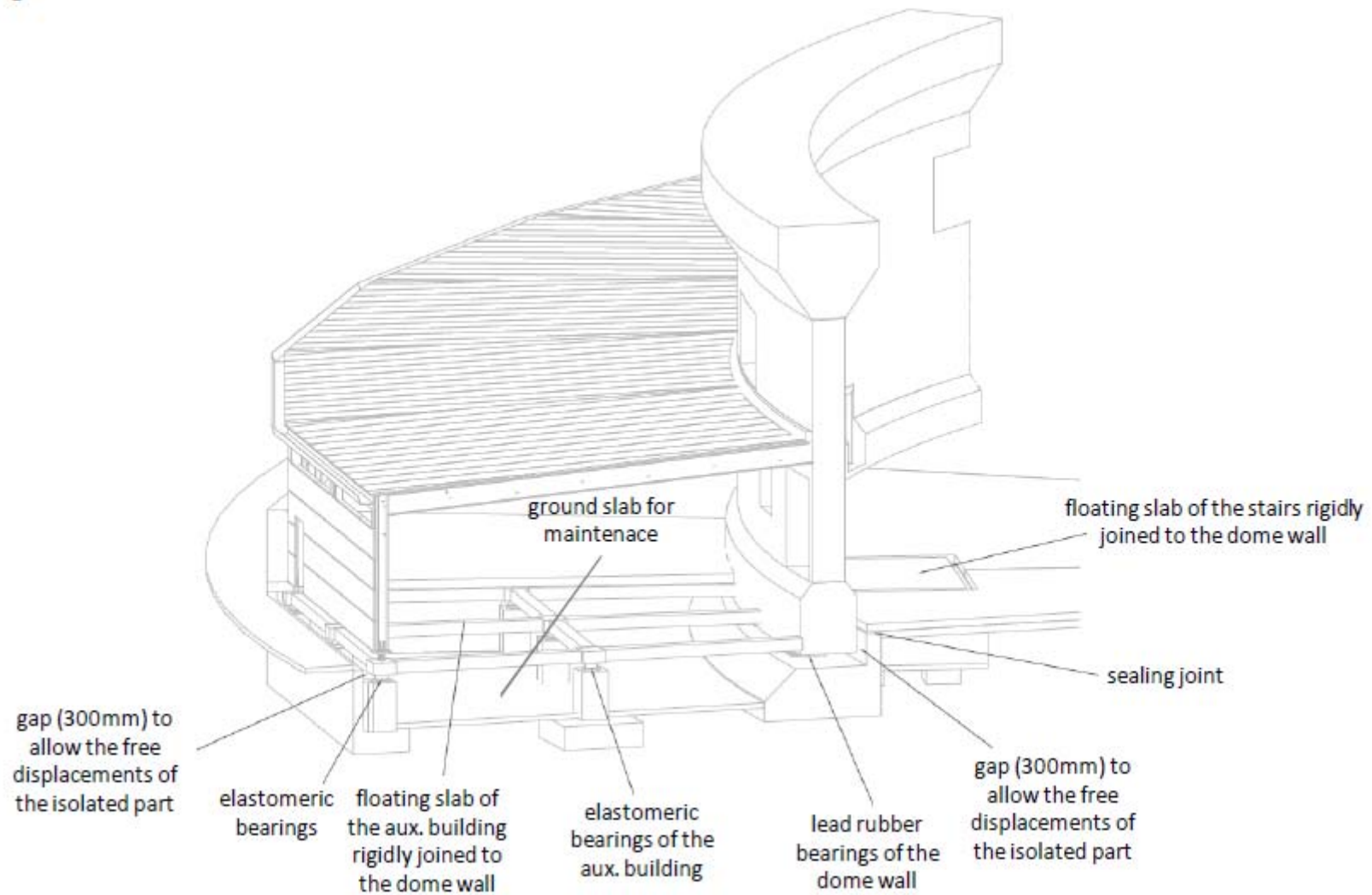
SUB-STRUCTURE



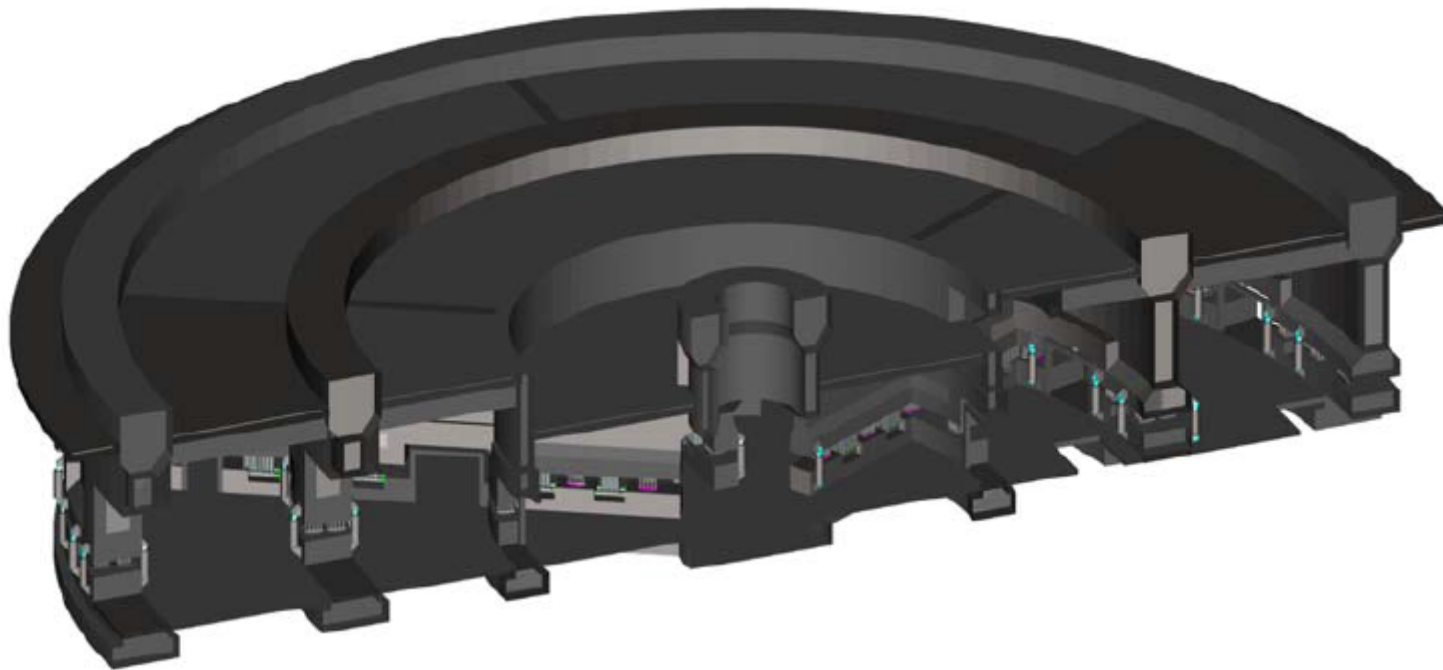
DOME FOUNDATION



CROSS SECTION



TELESCOPE FOUNDATION



SEISMIC ISOLATION SYSTEM



Common springs combined
with dashpots



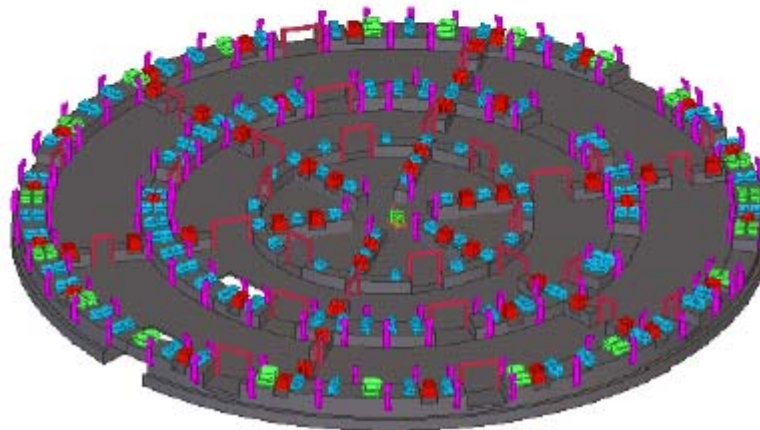
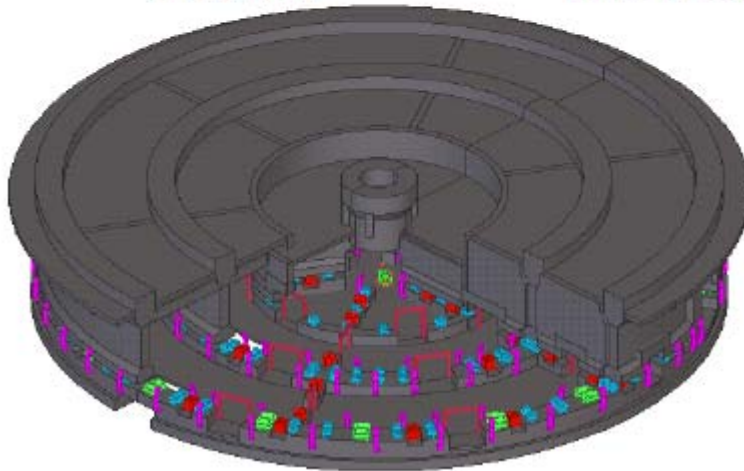
Common springs



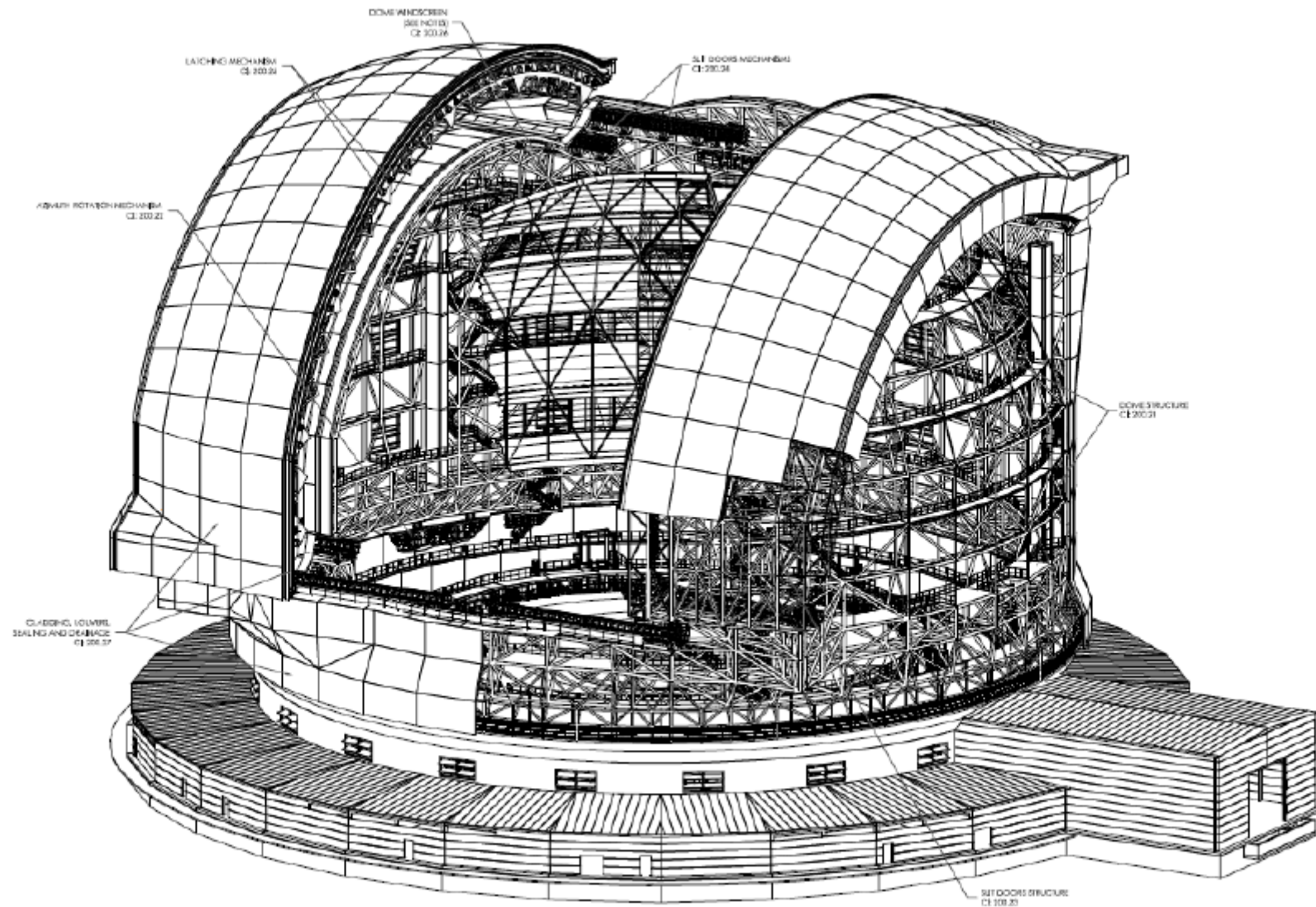
Vertical preloaded units



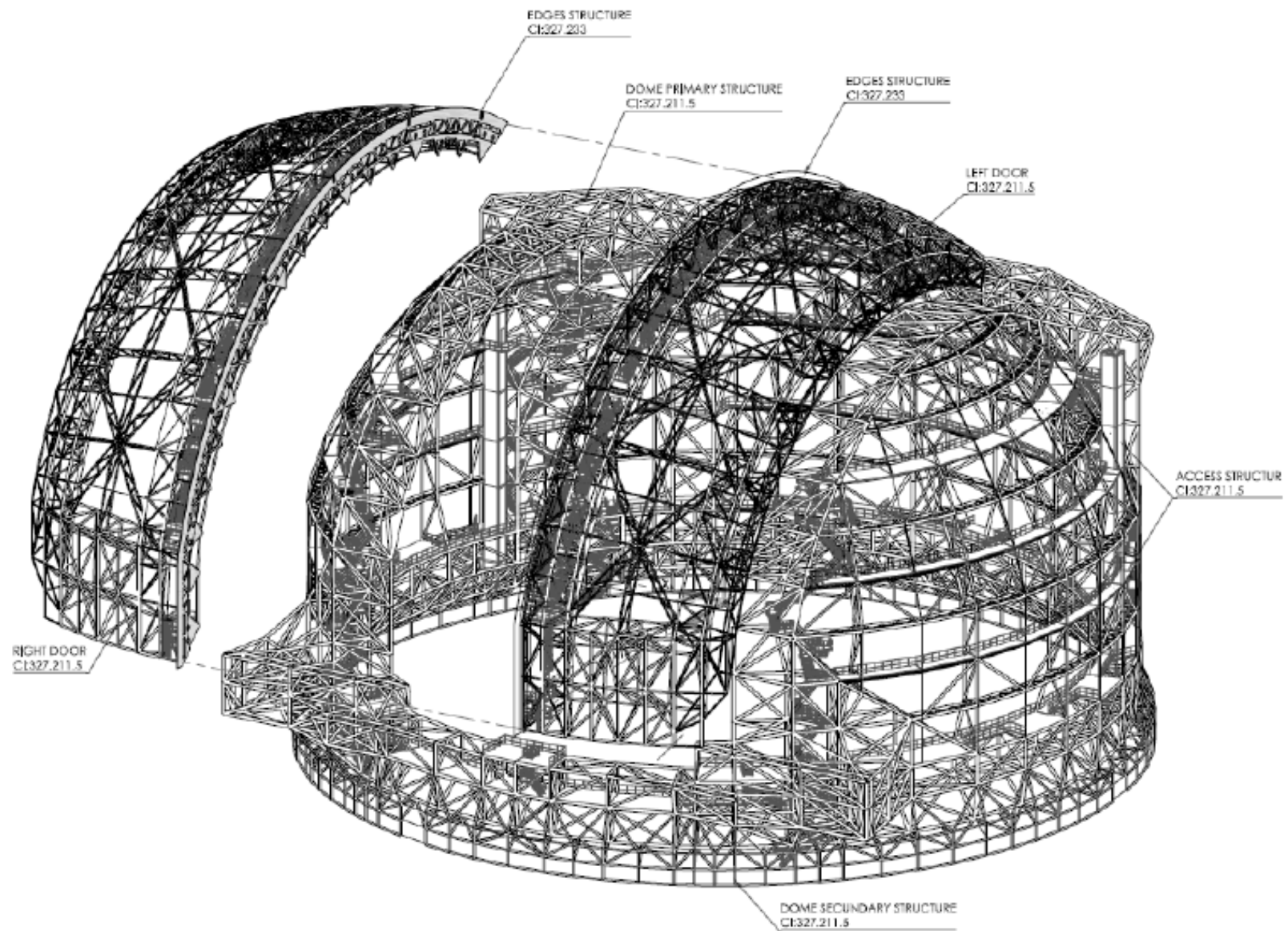
Horizontal preloaded units



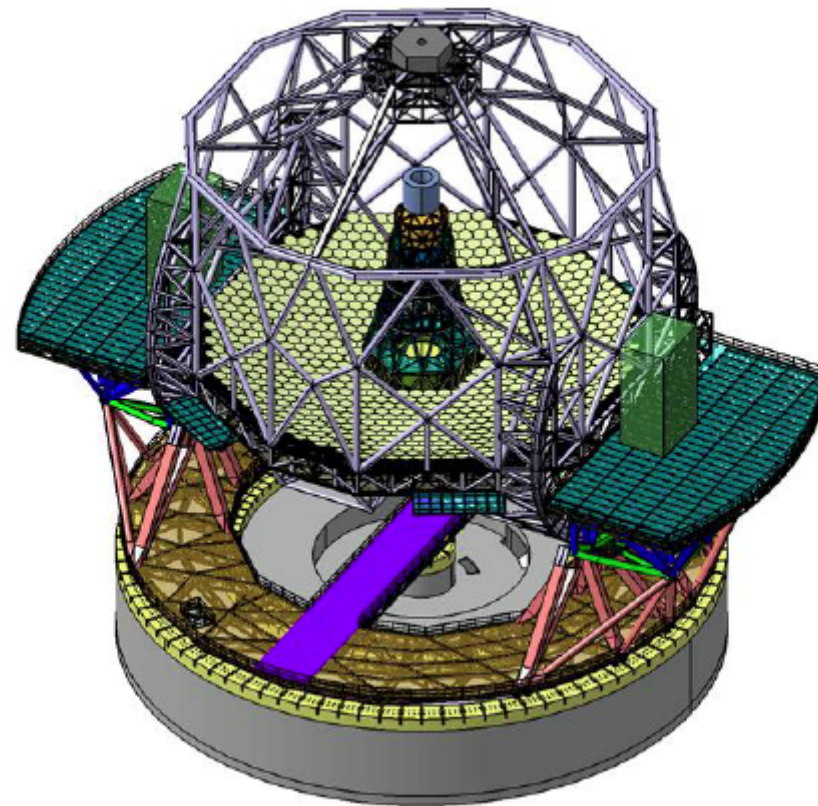
DOME STRUCTURE (WITHOUT TELESCOPE)



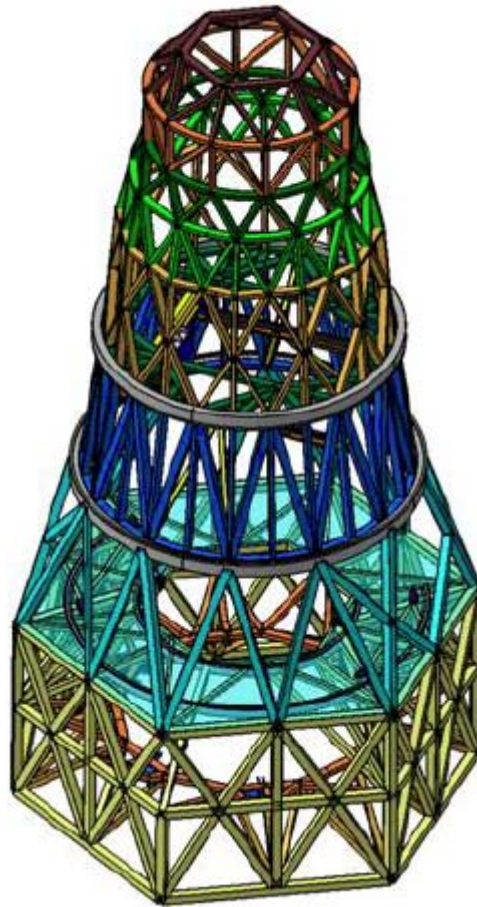
DOME STRUCTURE



MAIN TELESCOPE STRUCTURE



TELESCOPE STRUCTURE ADAPTIVE RELAY TOWER



RAMBOLL'S ROLE

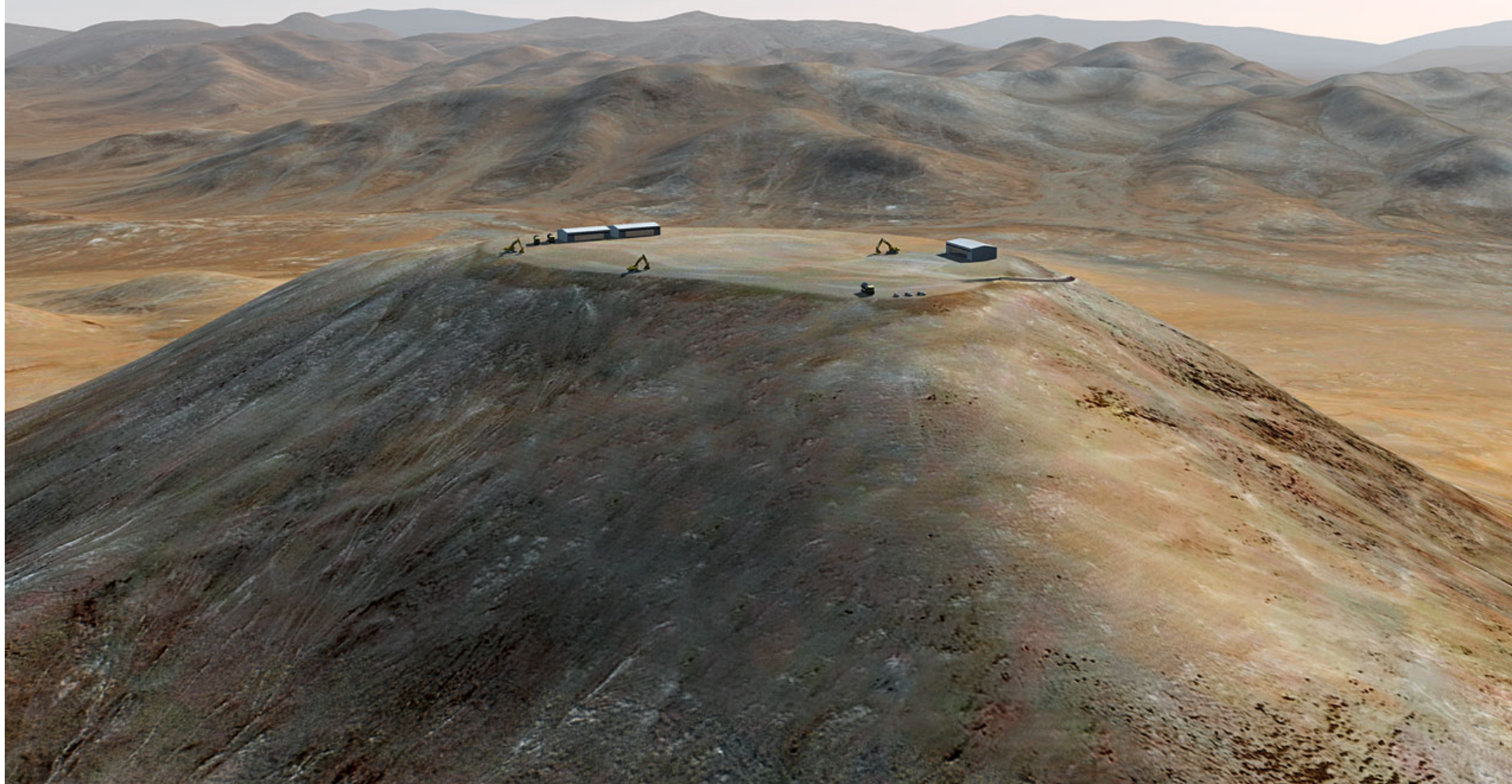
Client Consultant / Client's Engineer

- ESO has never built a telescope like this before
- ESO's in-house skills are mainly within technology
- There are a lot of uncertainties
- Getting the right experience involved is crucial
- An independent consultant was needed
- Ramboll will be involved in the next 10 years

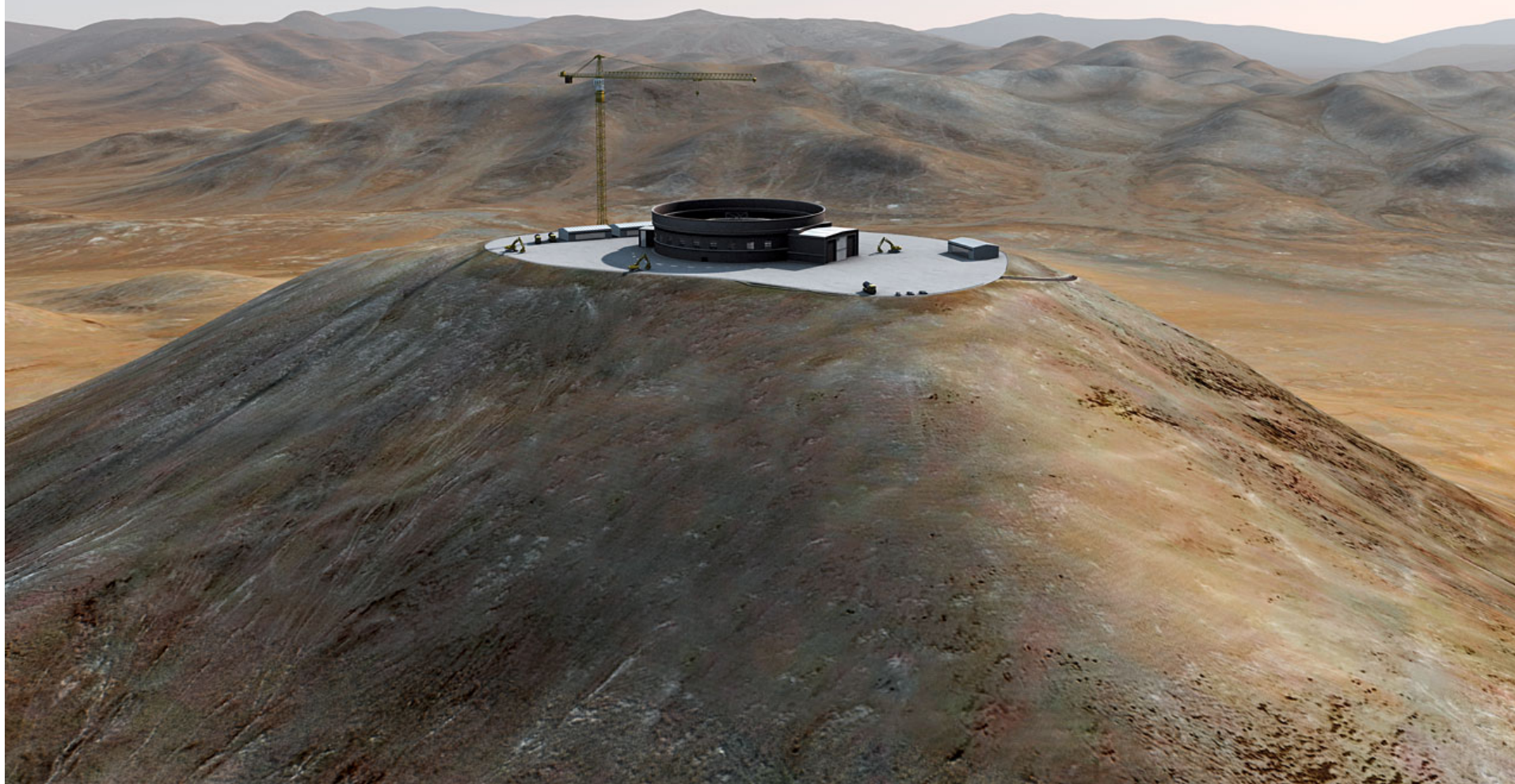
2014



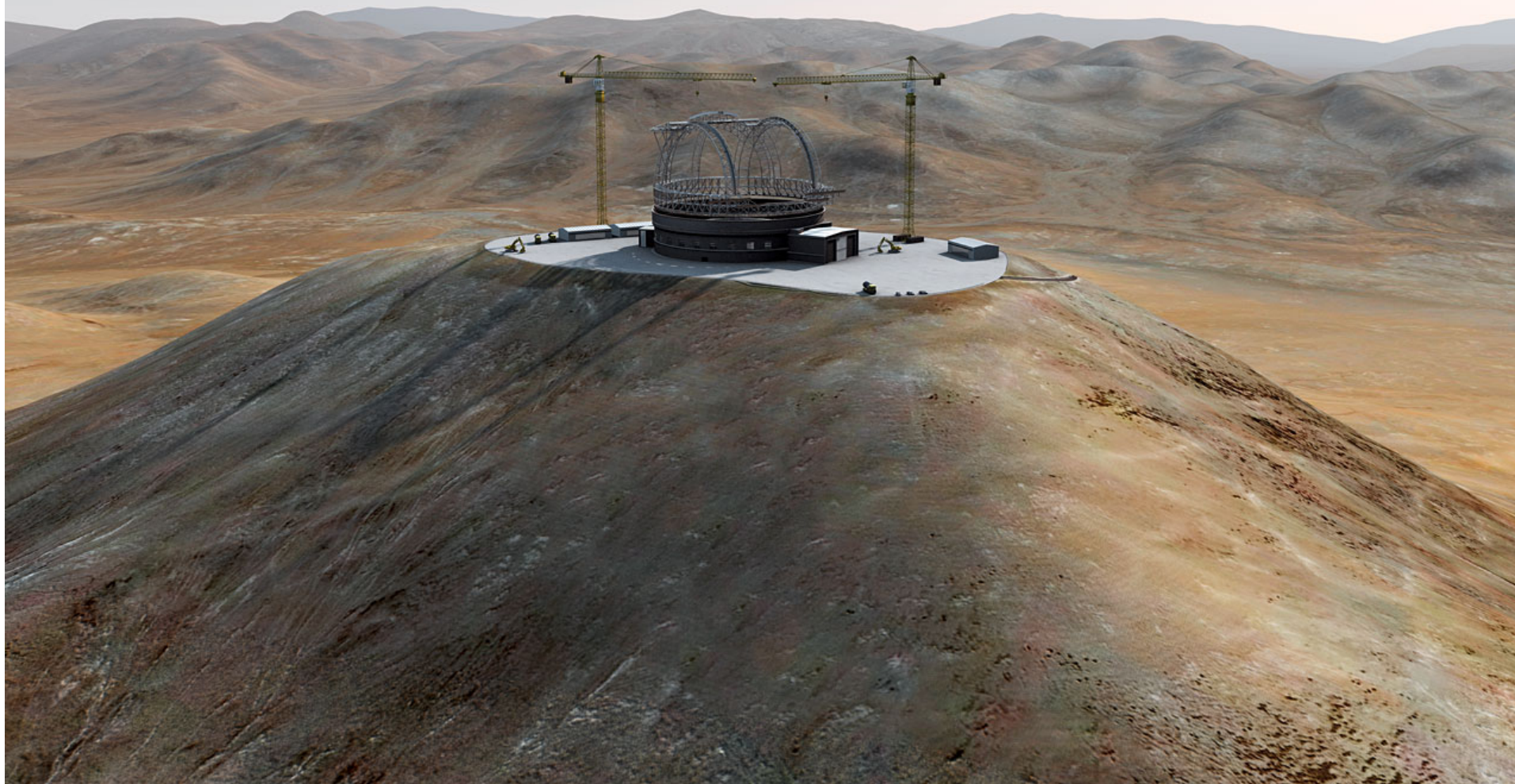
2015



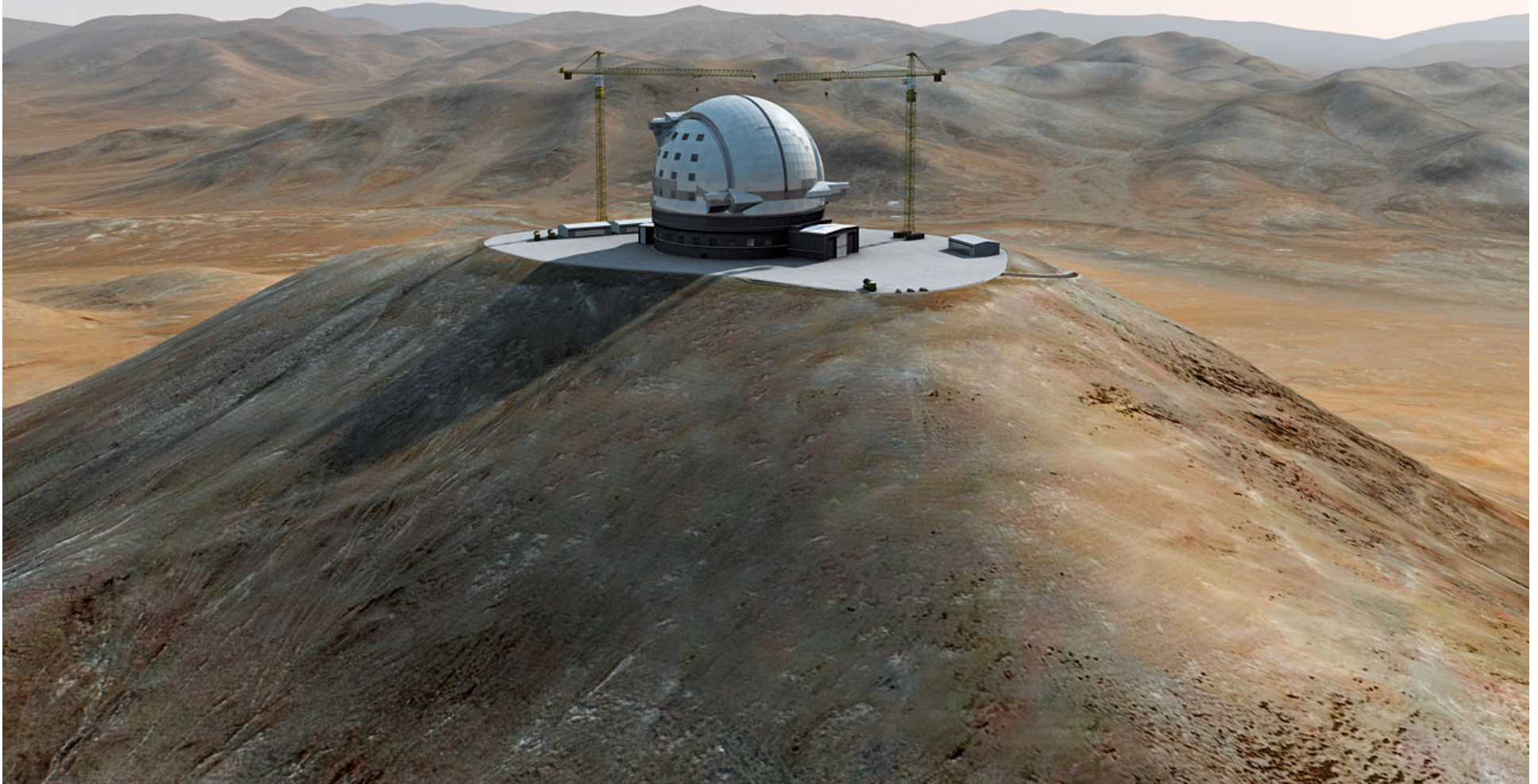
2016



2017



2018



2022

THE END (OF THE UNIVERSE)